

NRC Publications Archive Archives des publications du CNRC

Entanglement over global distances via quantum repeaters with satellite links

Boone, K.; Bourgoin, J.-P.; Meyer-Scott, E.; Heshami, K.; Jennewein, T.; Simon, C.

This publication could be one of several versions: author's original, accepted manuscript or the publisher's version. / La version de cette publication peut être l'une des suivantes : la version prépublication de l'auteur, la version acceptée du manuscrit ou la version de l'éditeur.

For the publisher's version, please access the DOI link below. / Pour consulter la version de l'éditeur, utilisez le lien DOI ci-dessous.

Publisher's version / Version de l'éditeur:

<https://doi.org/10.1103/PhysRevA.91.052325>

Physical Review A, 91, 5, 2015-05-26

NRC Publications Archive Record / Notice des Archives des publications du CNRC :

<https://nrc-publications.canada.ca/eng/view/object/?id=e53b8f87-3e49-4b6f-8e1b-ad2854e6c435>

<https://publications-cnrc.canada.ca/fra/voir/objet/?id=e53b8f87-3e49-4b6f-8e1b-ad2854e6c435>

Access and use of this website and the material on it are subject to the Terms and Conditions set forth at

<https://nrc-publications.canada.ca/eng/copyright>

READ THESE TERMS AND CONDITIONS CAREFULLY BEFORE USING THIS WEBSITE.

L'accès à ce site Web et l'utilisation de son contenu sont assujettis aux conditions présentées dans le site

<https://publications-cnrc.canada.ca/fra/droits>

LISEZ CES CONDITIONS ATTENTIVEMENT AVANT D'UTILISER CE SITE WEB.

Questions? Contact the NRC Publications Archive team at

PublicationsArchive-ArchivesPublications@nrc-cnrc.gc.ca. If you wish to email the authors directly, please see the first page of the publication for their contact information.

Vous avez des questions? Nous pouvons vous aider. Pour communiquer directement avec un auteur, consultez la première page de la revue dans laquelle son article a été publié afin de trouver ses coordonnées. Si vous n'arrivez pas à les repérer, communiquez avec nous à PublicationsArchive-ArchivesPublications@nrc-cnrc.gc.ca.



Entanglement over global distances via quantum repeaters with satellite links

K. Boone,¹ J.-P. Bourgoin,² E. Meyer-Scott,² K. Heshami,^{1,3,*} T. Jennewein,^{2,4,†} and C. Simon^{1,‡}

¹*Institute for Quantum Science and Technology and Department of Physics and Astronomy, University of Calgary, Calgary, Alberta, Canada T2N 1N4*

²*Institute for Quantum Computing, University of Waterloo, Waterloo, Ontario, Canada N2L 3G1*

³*National Research Council of Canada, 100 Sussex Drive, Ottawa, Ontario, Canada K1A 0R6*

⁴*Quantum Information Science Program, Canadian Institute for Advanced Research, Toronto, Ontario, Canada M5G 1Z8*

(Received 20 October 2014; published 26 May 2015)

We study entanglement creation over global distances based on a quantum repeater architecture that uses low-Earth-orbit satellites equipped with entangled photon sources, as well as ground stations equipped with quantum nondemolition detectors and quantum memories. We show that this approach allows entanglement creation at viable rates over distances that are inaccessible via direct transmission through optical fibers or even from very distant satellites.

DOI: [10.1103/PhysRevA.91.052325](https://doi.org/10.1103/PhysRevA.91.052325)

PACS number(s): 03.67.Hk, 03.65.Ud, 03.67.Dd, 42.50.Ex

Over the past few decades the distribution of quantum entanglement has progressed from tabletop experiments to distances of over 100 km [1]. Will it be possible to create entanglement over global distances? This is interesting from a fundamental point of view, but also from the perspective of trying to create a global quantum internet [2]. In the context of quantum cryptography, it would enable secure global communication without having to rely on any trusted nodes [3], as entanglement is the foundation for device-independent quantum key distribution [4]. It would also be useful for global clock networks [5] and for very long baseline telescopes [6].

Modern classical telecommunication relies on optical fibers. Unfortunately, the direct transmission of photons through fibers is not practical for quantum communication over global distances because losses are too high. The best available fibers have a loss of 0.15 dB/km at the optimal wavelength. This means, for example, that the time to distribute one entangled photon pair over 2000 km with a 1-GHz source exceeds the age of the universe.

Two alternative approaches to try to overcome this problem are currently being pursued in parallel, namely, fiber-based quantum repeaters and direct satellite links. Conventional quantum repeaters rely on first creating and storing entanglement for elementary links and then extending the distance of entanglement by entanglement swapping [7,8]. Based on the experimental and theoretical progress in this area over the past few years, it is plausible that this approach will make it possible to extend the distance of entanglement distribution significantly beyond what is possible with direct transmission through optical fibers [8–10]. However, truly global distances are still very difficult to envision for repeaters based on fiber links. This is true also for related approaches based on quantum error correction [11], which tend to require repeater stations that are only a few kilometers apart, such that global distances would imply thousands of repeater stations with hundreds of qubits per station.

The use of satellite links for quantum communication is also being pursued very actively. There has been a great deal of progress in terms of feasibility studies [12–19]. The launch of the first satellite carrying an entangled pair source has been announced for 2015 or 2016 [20]. The advantage of quantum communication via satellites is that transmission loss is dominated by diffraction rather than absorption and thus scales much more favorably with distance. For example, consider a pair source on a satellite at a height of 1000 km. For realistic assumptions (such as telescope size; see below), the combined transmission loss for the photon pair for a 2000-km ground station distance is only of order 40 dB. This should be contrasted with 300 dB for a fiber link of the same length. However, global distances are still challenging even for satellite links. Direct transmission from low-Earth-orbit (LEO) satellites, i.e., those below the Van Allen radiation belt, or up to about 2000 km in height, no longer works. Even before the Earth gets in the way, the loss becomes forbidding for very grazing incidence

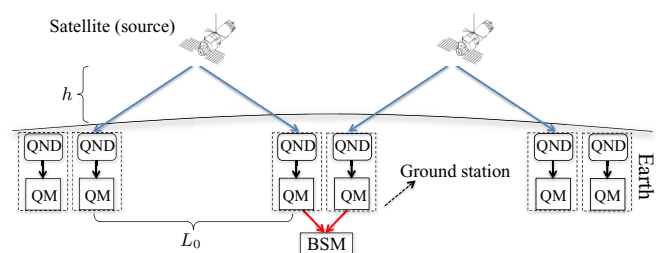


FIG. 1. (Color online) Proposed quantum repeater architecture with satellite links. Each elementary link (of length L_0) consists of an entangled photon pair source on a low-Earth-orbit satellite (at height h) and two ground stations consisting of quantum nondemolition (QND) measurement devices and quantum memories (QM). The successful transmission of entangled photons to each ground station is heralded by the QND devices, which detect the presence of a photon nondestructively and without revealing its quantum state. The entanglement is then stored in the memories until information about successful entanglement creation in two neighboring links is received. Then the entanglement can be extended by entanglement swapping based on a Bell state measurement (BSM). Figure 2 shows that four to eight such links are sufficient for spanning global distances.

*khabat.heshami@nrc.gc.ca

†thomas.jennewein@uwaterloo.ca

‡csimo@ucalgary.ca

due to the long propagation distance in air. One possible solution is to use satellites that are much further away, but this comes at significant cost, as satellites have to be much more robust to shield them from radiation. Moreover, the greatest ground distances, approaching 20 000 km (i.e., half the Earth's circumference), are out of range even for very distant satellites.

Here we propose to combine the two approaches discussed above. We study quantum repeaters based on LEO satellite links, as illustrated in Fig. 1. The satellites just need to be equipped with entangled pair sources, while the more complex components, such as quantum memories and quantum non-demolition (QND) detectors, are on the ground and can be further developed even after the satellites are launched. An important difference between satellite and fiber-based links is that the satellite-based links are active only during each time period when the satellite is visible from both ground stations (the flyby time T_{FB}). For currently realistic quantum memory lifetimes all satellite links in Fig. 1 have to be active simultaneously, which implies that our architecture requires a number of satellites equal to the number of links. However, our results show that four to eight links are sufficient to span global distances. For the present work we consider a simple situation where the stations are on the equator and the satellites are following each other around the equator. A true global network capable of linking arbitrary points across the globe would require a more complex configuration and a larger number of satellites. Let us note that there is a current trend in the space industry towards deploying large numbers of small and cheap LEO satellites, e.g., Planet Labs has recently deployed 71 Earth-imaging satellites [21].

Figure 2 compares the expected entanglement distribution rates per day for repeater architectures with LEO satellites to those achievable by direct transmission from more distant satellites. It is important to make the comparison on a per day basis since the flyby times and periods are different for satellites at different heights. Our results suggest that the approach based on repeaters with LEO satellite links is viable for all but the shortest distances and is the only way to create entanglement for the longest distances. We now describe the assumptions and requirements underlying these results in some detail.

One key ingredient for our analysis is the calculation of the probability for a pair of photons that are emitted from a satellite at height h to be successfully transmitted to the ground. Our approach, which is based on Ref. [17], takes into account diffraction, pointing error, and atmospheric transmittance. In Fig. 2 we assume a satellite transmitter size of 50 cm and an effective ground telescope size of 1 m. In practice it may be advantageous to use an array of smaller telescopes for the ground station to mitigate the effect of turbulence (see the Supplemental Material [22]). For the quantum repeater scenarios we assume a pair source that emits photons at 580 nm, which is motivated by our choice of quantum memory material (Eu-doped yttrium orthosilicate; see below). For direct transmission we assume a wavelength of 670 nm (470 nm) for $h = 2000$ km ($h = 10\,000$ km and geostationary), which results in optimal transmittance as shown in Ref. [17]. We include a satellite pointing error of $0.5 \mu\text{rad}$, which is an ambitious but realistic value [23], and assume ground stations at rural atmosphere at sea level; see the Supplemental Material [22] for more details. We assume that frequency shifts due to relativistic and gravitational effects [24] are compensated

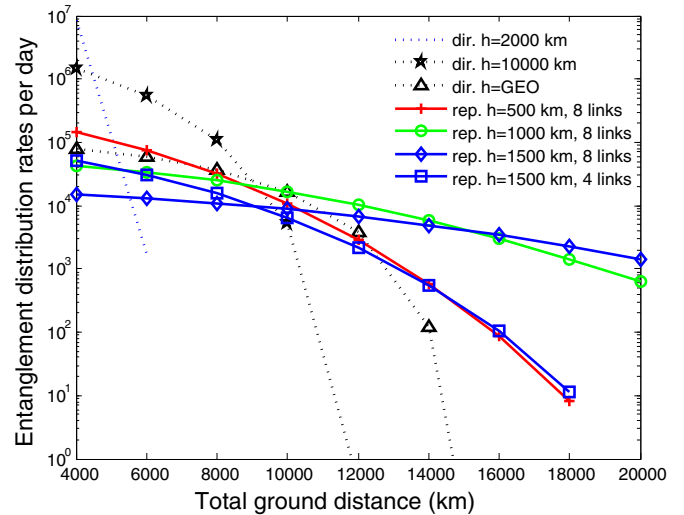


FIG. 2. (Color online) Rates of entangled pairs created per day as a function of ground distance for quantum repeaters with LEO satellites (solid lines) at heights $h = 500$, 1000 , and 1500 km, compared to direct transmission (dotted lines) from satellites at heights $h = 2000$ and $10\,000$ km and from a geostationary satellite.

(e.g., by acousto-optic modulators on the ground). Timing jitter due to turbulence in the atmosphere is negligible for the relatively long pulses that we are considering [25].

For the repeater scenarios, we have assumed a pair source with a repetition rate of 10 MHz. This value is motivated primarily by the expected memory bandwidth for our choice of material; see below. In contrast, we assume a much higher 1-GHz repetition rate for direct transmission. In each case the source could, e.g., be a deterministic pair source based on a quantum dot in microcavity [26]. However, simpler implementations are possible based on parametric down-conversion sources with a small pair creation probability per pulse (below 0.01) [27], in order to avoid errors due to multipair emissions. If one aims to achieve the same effective rate in this way, the underlying repetition rate (and hence memory bandwidth, in the repeater scenario) has to be increased correspondingly. Memory bandwidths up to 1 GHz have already been achieved in rare-earth-doped materials (e.g., in Tm:LiNbO [28]), but not yet in combination with long storage times. We have not assumed any frequency multiplexing either for repeaters or for direct transmission. This could be used to boost rates in both scenarios, at the expense of more complex sources on the satellites.

The rates for the repeaters are calculated as in Ref. [8], assuming a nested approach. That is, entanglement is first created and stored at the level of the elementary links. Then links are connected in a hierarchical fashion, forming links of length two, four, etc. For convenience let us define the average probability of a pair reaching the ground stations during one flyby of the satellite as $P_0^{\text{av}} = \int \eta_{\text{tr}}^{(2)}(t) dt / T_{\text{FB}}$, where $\eta_{\text{tr}}^{(2)}(t)$ is the probability for both photons to be transmitted from the satellite to the ground stations at a given time t and T_{FB} is the flyby time of the satellite [22]. The probability of successfully creating, transmitting, and storing an entangled pair over one elementary link is $P_{\text{EG}} = \eta_s P_0^{\text{av}} \eta_f^2 \eta_q^2 \eta_w^2$, where η_s , η_q , and η_w are source, QND detector, and memory write efficiencies.

Entanglement swapping relies on Bell-state measurements (BSMs). In our scheme, a successful BSM requires successful readout of two photons from neighboring quantum memories with the efficiency of η_r^2 and two single-photon detections with η_d^2 efficiency. Here η_r and η_d are memory readout and detector efficiencies. This gives the entanglement swapping efficiency of $P_{\text{ES}} = \frac{\eta_r^2 \eta_d^2}{2}$, where the factor of $\frac{1}{2}$ is due to limited success probability of the BSM using linear optics with ancillary vacuum modes [29]. Higher success probabilities are possible in principle using ancillary photons [30,31]. For a repeater composed of 2^n links, the number of entangled pairs created during one flyby is given by $R_s T_{\text{FB}} P_{\text{EG}} (\frac{2}{3} P_{\text{ES}})^n$, where R_s is the source rate. The factors of $\frac{2}{3}$ take into account the fact that entanglement has to be created in two neighboring links before entanglement swapping can proceed [8]. In Fig. 2 we assume $\eta_s = \eta_w = \eta_r = \eta_d = 0.9$, which are ambitious, but realistic numbers given the current state of technology. In contrast, we only assume $\eta_q = 0.32$, taking into account the fact that the QND detection is likely to require coupling into a single-mode waveguide, which is difficult to do perfectly in the presence of turbulence (see also the Supplemental Material [22]). The impact of changing these efficiencies on the total entanglement distribution rate is shown in Fig. 3 and in the Supplemental Material [22].

Quantum memories for photons have been implemented in a range of physical systems [32]. Memories based on rare-earth ion-doped crystals [33] are particularly attractive for our purpose because of their potential for highly multimode storage, e.g., using the atomic frequency comb (AFC) protocol [34]. This is important because the quantum memories in each ground station will be exposed to a large number of photons $N_{\text{mod}} = R_s \eta_s \eta_{\text{tr,max}}^{(1)} \frac{L_0}{c}$ before receiving the classical signals from the other end of each link that make it possible to decide which photons are part of an entangled pair and should thus be kept for entanglement swapping. Here $\eta_{\text{tr,max}}^{(1)}$ denotes the maximum value of the single-photon transmission during one flyby. For the quantum repeater scenarios in Fig. 2, multimode storage of up to several thousand photons is required according to the above formula (depending on satellite height and number of links). A single AFC-type quantum memory based on Eu-doped yttrium ortho silicate (YSO) should be able to store 10^2 – 10^3 photons in distinct temporal modes [34]. Having several waveguides or using multiple locations on the same crystal makes the storage of thousands of photons in distinct modes in a single crystal plausible. Our protocol also requires storage times of the order of the total communication time L/c , where L is the total distance, which corresponds to 67 ms for 20 000 km. Such long storage times should be achievable by transferring the optical memory excitations to ground spin states [35,36]. Reference [37] recently demonstrated spin coherence times of several hours in Eu-doped YSO. The requirement of transferring the excitation to the ground state limits the repetition rate of the photon source as the bandwidth of the photons must be smaller than energy spacing between the ground spin states. The 10-MHz bandwidth assumed in Fig. 2 is compatible with the ground level separations of Eu:YSO, which are of order 100 MHz [36]. High memory efficiencies can be achieved in rare-earth-doped crystals with the help of optical cavities [38].

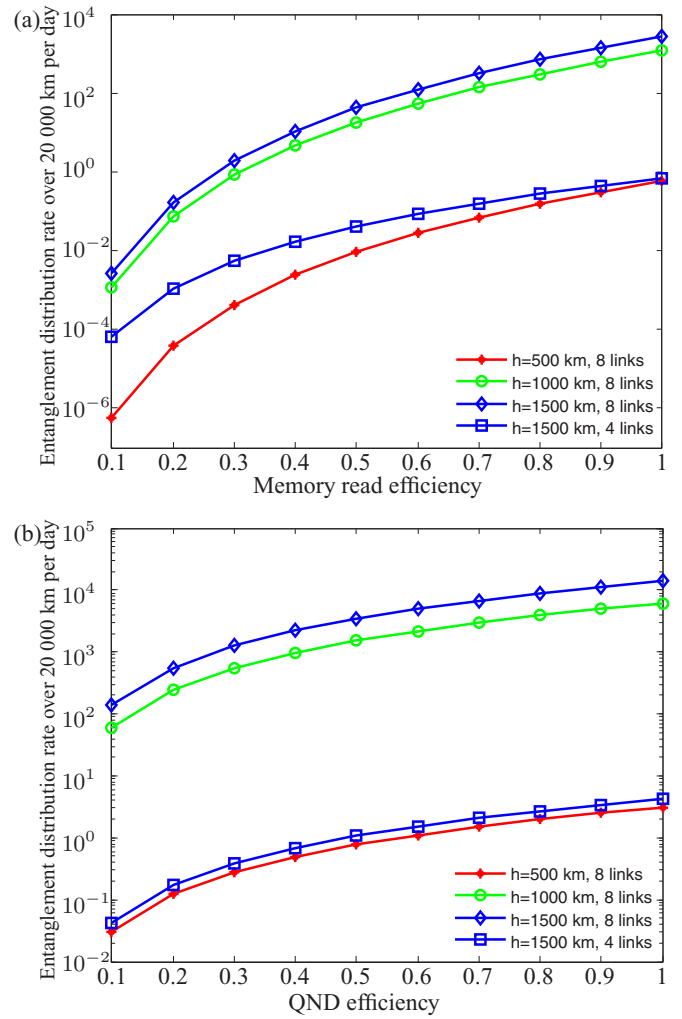


FIG. 3. (Color online) Impact of inefficiencies in various elements on the entanglement distribution rate over 20 000 km for the repeater protocols shown in Fig. 2. (a) Effect of the memory read efficiency. The detector efficiency has the same effect. (b) Effect of QND detector efficiency. Memory write efficiency and source efficiency have similar effects (see also the Supplemental Material [22]). The repeater protocol is more sensitive to memory read and detector efficiency than to that of the other components, because the former efficiencies intervene in each entanglement swapping step, whereas the latter only intervene in the entanglement creation in the elementary links.

Our scheme also requires QND detection of the photonic qubits. Quantum nondemolition measurement of photons has recently been demonstrated using a single atom in a cavity [39]. The cross Kerr effect induced by the ac Stark shift in atomic ensembles also provides the possibility to realize QND measurement of photons. In Ref. [40], 0.5-mrad cross-phase shift per photon has been shown using a hot atomic vapor inside a hollow-core photonic crystal fiber, which should already allow a QND measurement of the photon number [41]. Here we also require the QND measurement to be insensitive to the photonic qubit state. For example, if photon pairs with polarization entanglement are to be detected, the probe field must interact with both polarization modes. A simpler

implementation of the QND detection of photonic qubits is possible for time-bin qubits based on the ac Stark shift in combination with quantum storage because the phase shift imparted to the stored probe beam is not sensitive to the precise timing of the signal photon propagating through the ensemble [42]. This approach should also make it possible to integrate the QND detector with the quantum memory, e.g., a rare-earth-doped waveguide [28]. Another possibility is to use a heralded qubit amplifier based on linear optics and a deterministic pair source [43]. This achieves a QND detection efficiency of up to 0.5.

We only performed a simple rate calculation for the proposed repeater architecture. A more sophisticated analysis would characterize the fidelity of the distributed quantum state and extract a key rate for quantum key distribution applications [10]. However, assuming low noise levels in all components and given the fact that we only consider

small numbers of repeater links, the present estimates should give a reasonably accurate picture of achievable key rates.

We have argued that quantum repeaters based on LEO satellite links are a viable approach to global quantum communication. Our proposed scheme relies on realistic advances in quantum memories and quantum nondemolition measurements and only requires a moderate number of satellites equipped with entangled photon pair sources. Ultimately, global quantum repeater networks will likely combine satellite links for very long distances with fiber links for short and intermediate distances.

The authors would like to thank W. Tittel for valuable discussions. This work was supported by NSERC, AITF, DARPA, the Ontario Ministry of Research and Innovation, ERA, Industry Canada, CIFAR, CFI, and FEDDEV Ontario.

-
- [1] S. J. Freedman and J. F. Clauser, *Phys. Rev. Lett.* **28**, 938 (1972); A. Aspect, J. Dalibard, and G. Roger, *ibid.* **49**, 1804 (1982); G. Weihs, T. Jennewein, C. Simon, H. Weinfurter, and A. Zeilinger, *ibid.* **81**, 5039 (1998); W. Tittel, J. Brendel, H. Zbinden, and N. Gisin, *ibid.* **81**, 3563 (1998); R. Ursin *et al.*, *Nat. Phys.* **3**, 481 (2007); X.-S. Ma *et al.*, *Nature (London)* **489**, 269 (2012).
- [2] H. J. Kimble, *Nature (London)* **453**, 1023 (2008).
- [3] T. Jennewein and B. Higgins, *Phys. World* **26**(3), 52 (2013).
- [4] A. K. Ekert, *Phys. Rev. Lett.* **67**, 661 (1991); A. Acin, N. Brunner, N. Gisin, S. Massar, S. Pironio, and V. Scarani, *ibid.* **98**, 230501 (2007).
- [5] P. Kómár, E. M. Kessler, M. Bishof, L. Jiang, A. S. Sørensen, J. Ye, and M. D. Lukin, *Nat. Phys.* **10**, 582 (2014).
- [6] D. Gottesman, T. Jennewein, and S. Croke, *Phys. Rev. Lett.* **109**, 070503 (2012).
- [7] H.-J. Briegel, W. Dür, J. I. Cirac, and P. Zoller, *Phys. Rev. Lett.* **81**, 5932 (1998); L.-M. Duan, M. D. Lukin, J. I. Cirac, and P. Zoller, *Nature (London)* **414**, 413 (2001).
- [8] N. Sangouard, C. Simon, H. de Riedmatten, and N. Gisin, *Rev. Mod. Phys.* **83**, 33 (2011).
- [9] F. Bussières, N. Sangouard, M. Afzelius, H. de Riedmatten, C. Simon, and W. Tittel, *J. Mod. Opt.* **60**, 1519 (2013); N. Sinclair, E. Saglamyurek, H. Mallahzadeh, J. A. Slater, M. George, R. Ricken, M. P. Hedges, D. Oblak, C. Simon, W. Sohler, and W. Tittel, *Phys. Rev. Lett.* **113**, 053603 (2014).
- [10] S. Guha, H. Krovi, C. A. Fuchs, Z. Dutton, J. A. Slater, C. Simon, and W. Tittel, [arXiv:1404.7183](https://arxiv.org/abs/1404.7183).
- [11] W. J. Munro, A. M. Stephens, S. J. Devitt, K. A. Harrison, and K. Nemoto, *Nat. Photon.* **6**, 777 (2012); K. Azuma, K. Tamaki, and H.-K. Lo, *Nat. Commun.* **6**, 6787 (2015); S. Muralidharan, J. Kim, N. Lütkenhaus, M. D. Lukin, and L. Jiang, *Phys. Rev. Lett.* **112**, 250501 (2014).
- [12] W. T. Buttler, R. J. Hughes, P. G. Kwiat, S. K. Lamoreaux, G. G. Luther, G. L. Morgan, J. E. Nordholt, C. G. Peterson, and C. M. Simmons, *Phys. Rev. Lett.* **81**, 3283 (1998).
- [13] C. Kurtsiefer, P. Zarda, M. Halder, H. Weinfurter, P. M. Gorman, P. R. Tapster, and J. G. Rarity, *Nature (London)* **419**, 450 (2002).
- [14] J. G. Rarity, P. R. Tapster, P. M. Gorman, and P. Knight, *New J. Phys.* **4**, 82 (2002).
- [15] C. Bonato, A. Tomaello, V. Da Deppo, G. Naletto, and P. Villoresi, *New J. Phys.* **11**, 045017 (2009).
- [16] J.-Y. Wang *et al.*, *Nat. Photon.* **7**, 387 (2013).
- [17] J.-P. Bourgoin, E. Meyer-Scott, B. L. Higgins, B. Helou, C. Erven, H. Hübel, B. Kumar, D. Hudson, I. D'Souza, R. Girard, R. Laflamme, and T. Jennewein, *New J. Phys.* **15**, 023006 (2013).
- [18] Z. Tang, R. Chandrasekara, Y. Y. Sean, C. Cheng, C. Wildfeuer, and A. Ling, *Sci. Rep.* **4**, 6366 (2014).
- [19] G. Vallone, D. Bacco, D. Dequal, S. Gaiarin, V. Luceri, G. Bianco, and P. Villoresi, [arXiv:1406.4051](https://arxiv.org/abs/1406.4051).
- [20] N. Horiuchi, *Nat. Photon.* **9**, 13 (2015).
- [21] <https://www.planet.com/story/>.
- [22] See Supplemental Material at <http://link.aps.org/supplemental/10.1103/PhysRevA.91.052325> for more details on the transmission calculation, the effect of turbulence on single-mode fiber coupling, the satellite orbits, the effects of inefficiencies on entanglement distribution rates, the errors due to background light, and a potential repeater architecture based on uplinks rather than downlinks.
- [23] M. Toyoda, *Opt. Eng.* **45**, 036001 (2006).
- [24] D. E. Bruschi, T. Ralph, I. Fuentes, T. Jennewein, and M. Razavi, *Phys. Rev. D* **90**, 045041 (2014).
- [25] L. Kral, I. Prochazka, and K. Hamal, *Opt. Lett.* **30**, 1767 (2005).
- [26] A. Dousse, J. Suffczynski, A. Beveratos, O. Krebs, A. Lemaitre, I. Sagnes, J. Bloch, P. Voisin, and P. Senellart, *Nature (London)* **466**, 217 (2010).
- [27] J. A. Slater, J.-S. Corbeil, S. Virally, F. Bussières, A. Kudlinski, G. Bouwmans, S. Lacroix, N. Godbout, and W. Tittel, *Opt. Lett.* **35**, 499 (2010).
- [28] E. Saglamyurek, N. Sinclair, J. Jin, J. A. Slater, D. Oblak, F. Bussières, M. George, R. Ricken, W. Sohler, and W. Tittel, *Nature (London)* **469**, 512 (2011).
- [29] J. Calsamiglia and N. Lütkenhaus, *Appl. Phys. B* **72**, 67 (2001).
- [30] W. P. Grice, *Phys. Rev. A* **84**, 042331 (2011).
- [31] F. Ewert and P. van Loock, *Phys. Rev. Lett.* **113**, 140403 (2014).
- [32] K. Hammerer, A. S. Sørensen, and E. S. Polzik, *Rev. Mod. Phys.* **82**, 1041 (2010); A. I. Lvovsky, B. C. Sanders, and W. Tittel,

- [Nat. Photon.](#) **3**, 706 (2009); C. Simon *et al.*, [Eur. Phys. J. D](#) **58**, 1 (2010).
- [33] H. de Riedmatten, M. Afzelius, M. U. Staudt, C. Simon, and N. Gisin, [Nature \(London\)](#) **456**, 773 (2008); M. P. Hedges, J. J. Longdell, Y. Li, and M. J. Sellars, *ibid.* **465**, 1052 (2010); W. Tittel, M. Afzelius, T. Chaneliere, R. L. Cone, S. Kroll, S. A. Moiseev, and M. Sellars, [Laser Photon. Rev.](#) **4**, 244 (2010).
- [34] M. Afzelius, C. Simon, H. de Riedmatten, and N. Gisin, [Phys. Rev. A](#) **79**, 052329 (2009).
- [35] M. Afzelius, I. Usmani, A. Amari, B. Lauritzen, A. Walther, C. Simon, N. Sangouard, J. Min'ar, H. de Riedmatten, N. Gisin, and S. Kröll, [Phys. Rev. Lett.](#) **104**, 040503 (2010).
- [36] P. Jobez *et al.*, [arXiv:1501.03981](#).
- [37] M. Zhong *et al.*, [Nature \(London\)](#) **517**, 177 (2015).
- [38] M. Afzelius and C. Simon, [Phys. Rev. A](#) **82**, 022310 (2010); M. Sabooni, Q. Li, S. Kröll, and L. Rippe, [Phys. Rev. Lett.](#) **110**, 133604 (2013).
- [39] A. Reiserer, N. Kalb, G. Rempe, and S. Ritter, [Nature \(London\)](#) **508**, 237 (2014).
- [40] V. Venkataraman, K. Saha, and A. L. Gaeta, [Nat. Photon.](#) **7**, 138 (2013).
- [41] N. Imoto, H. A. Haus, and Y. Yamamoto, [Phys. Rev. A](#) **32**, 2287 (1985).
- [42] K. Heshami *et al.* (unpublished).
- [43] M. Curty and T. Moroder, [Phys. Rev. A](#) **84**, 010304 (2011); E. Meyer-Scott, M. Bula, K. Bartkiewicz, A. Černoč, J. Soubusta, T. Jennewein, and K. Lemr, *ibid.* **88**, 012327 (2013).